Shell evolution and isomers below ¹³²Sn: Spectroscopy of neutron-rich ₄₆Pd and ₄₇Ag isotopes

<u>Hiroshi Watanabe</u> School of Physics, Beihang University, Beijing, China

The shell structures of atomic nuclei are nowadays known to change with the variation of the proton or neutron number, due predominantly to the monopole part of the proton-neutron interaction that includes the central and tensor forces [1]. Such a shell evolutionary behavior is expected to become pronounced when the proton-neutron imbalance is very large, leading to lost or new magic numbers [2]: For example, the conventional magic numbers N = 8, 20, and 28 disappear and the new magicity emerges at N = 16, 32, and 34, depending on the location of the nucleus in the N-Z plane. However, we don't know yet whether similar change of the shell structure can take place at the heavier conventional magic numbers N = 50, 82, and 126, which also play an important role in determining the solar abundance distribution particularly around the three prominent peaks at A \approx 80, 130, and 195, respectively, that would result from the rapid neutron-capture (r) process.

The neutron-rich isotopes of Pd (Z = 46) and Ag (Z = 47) have attracted considerable interest in terms of the evolution of the N = 82 shell closure and its influence on the r-process nucleosynthesis. Such previously unreachable exotic nuclides have become accessible by means of in-flight fission of a high-intensity ²³⁸U beam available at a new-generation RIbeam facility, the RI-Beam Factory (RIBF) in RIKEN Nishina Center [3]. In this presentation, recent spectroscopic results of Pd and Ag isotopes obtained as part of the EURICA (EUROBALL-RIKEN Cluster Array) project at RIBF [4] will be presented, with a particular focus on characteristic isomers, such as a seniority isomer in ¹²⁸Pd₈₂ [5], long-lived high-spin isomers in ¹²⁶Pd₈₀ [6] and ¹²⁷Ag₈₀, isomers with proton-hole and neutron-hole excitations in ^{125,127}Pd_{79,81} [7], and low-lying β -emitting isomers in ^{123,125}Ag_{76,78} [8]. The nature of these isomers will be discussed in terms of the effect of proton-neutron interactions and the resultant shell evolution below the doubly magic nucleus ¹³²Sn in the framework of shellmodel approaches.

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