

# Realistic shell model and nuclei around $^{132}\text{Sn}$

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In the last ten years or so, nuclei in the mass region around  $^{132}\text{Sn}$  have become accessible to experimental studies thanks to new radioactive ion beam facilities and the development of sophisticated detection techniques. These nuclei represent a crucial opportunity to test the main ingredients of the nuclear-shell model and investigate the evolution of the shell structure when going far from stability valley in heavy-mass nuclei.

In the light- and medium-mass regions, structural changes have been evidenced for nuclei with a large excess of neutrons, leading to the breakdown of the traditional magic numbers and the appearance of new ones. These findings have driven a great theoretical effort to understand the microscopic mechanism underlying the shell evolution, with special attention to the role of the different components of the nuclear force (see, for instance, [1]).

The available experimental data for nuclei around  $^{132}\text{Sn}$ , which are, however still scarce especially for systems with  $N > 82$ , have shown peculiar properties although no clear signatures of modifications in the shell structure.

In this contribution, I shall focus on some selected results for nuclei with a few valence particles and/or holes with respect to  $^{132}\text{Sn}$ , that have been obtained within the shell-model framework by using a microscopic effective interaction [2].

Calculations have been carried out by assuming a closed  $^{132}\text{Sn}$  core and including the  $0g_{9/2}1d_{5/2}0h_{11/2}$  and  $0h_{9/2}1f_{7/2}0i_{13/2}$  orbitals for proton particles/neutron holes and neutron particles, respectively. A unique shell-model Hamiltonian is adopted, with the single-particle(hole) energies taken from experiment and the two-body effective interaction derived by means of the many-body perturbation theory [3] from the CD-Bonn nucleon-nucleon potential [4] renormalized by means of the  $V_{\text{low-k}}$  approach [5].

Results are compared with experiments, and predictions that may provide guidance to future experiments are also discussed.

[1] T. Otsuka, A. Gade, O. Sorlin, T. Suzuki, Y. Utsuno, arXiv:1805.06501.

[2] L. Coraggio, A. Covello, A. Gargano, and N. Itaco, Phys. Rev. C **90**, 044322 (2014), and references therein.

[3] L. Coraggio, A. Covello, A. Gargano, N. Itaco, and T. T. S. Kuo, Prog. Part. Nucl. Phys. **62**, 135 (2009).

[4] R. Machleidt, Phys. Rev. C **63**, 024001 (2001).

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